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Dark quark nuggets 1810.04360

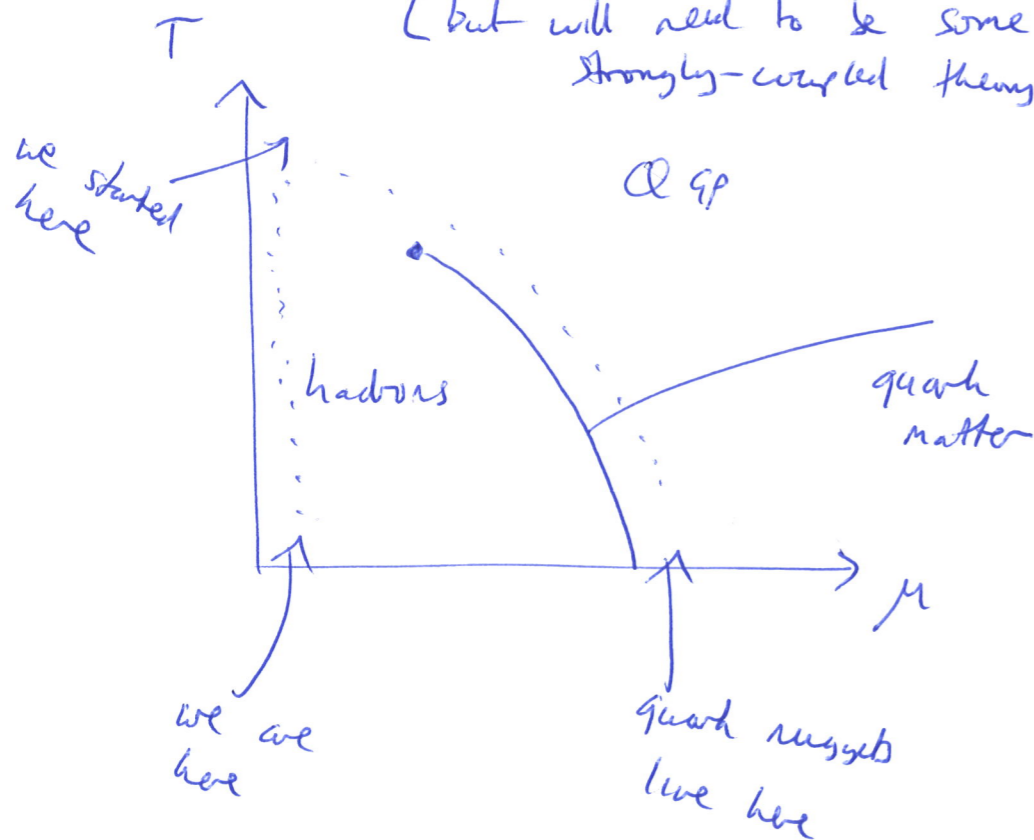
Inspired by Witten's 1984

"Cosmic separation of phases"

Quark nuggets:

- Macroscopic nucleons, $N_B > 10^{30}$, $r \lesssim 1 \text{ cm}$
- Formed when (most of) the universe undergoes FOPT (XSB + confinement)
- [not viable in SM]
- Could survive as DM candidate.

[but will need to be some other strongly-coupled theory besides QCD]



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What do nuggets need?

- 1) First order PT Q_1 Q_2
- 2) Conserved global charge, like baryon number
- 3) Cosmological excess of matter over antimatter.

↓
quark nuggets form & stick around

How?

- ① Possible in (e.g.) dark QCD and similar theories with $N_f \geq 3$ light fermions (QCD doesn't work because strange isn't very light)
- ② Dark baryon # (or technibaryon/twin baryon/etc)
- ③ Assume that visible baryon asymmetry is n_B/n_γ shared with dark sector (assume same μ)

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Dark QCD

$$\left. \begin{array}{l} N_d \text{ colours} \\ N_f \text{ light flavours} \end{array} \right\} \rightarrow \begin{array}{l} N_d^2 - 1 \text{ dark gluons} \\ N_f \text{ dark } q/\bar{q} \end{array}$$

Need a theory with a dark colour confining phase transition + χ SB



1/ quark masses are light ($m_i \ll \Lambda_d$) and $N_f \geq 3$, "Pisarski-Wilczek" argument says get FORT $\frac{2}{3}$ (with χ SB)

2/ $N_f \leq 2N_d + 4$ get confinement (it need to be below the confinement window)

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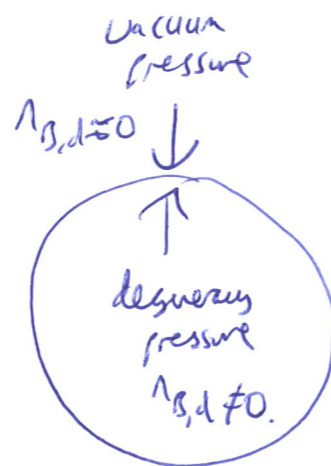
The $T=0$, high density region possesses a dark quark matter phase.

Assume temperature T , thermal eqⁿ

$$\mu \gg 0 \quad \Lambda_{B,d}$$

$$m_i \ll T \ll \mu$$

relativistic degenerate Fermi gas.



$$E_{\text{dark}} = 4B \quad \text{energy density}$$

$$\Lambda_{B,d,dQM} = \left(\frac{64 N_f}{3\pi^2 N_d^3} \right)^{1/4} B^{3/4} \quad \text{density of dark BQ}$$

$$\frac{E_{\text{dark}}}{\Lambda_{B,d,dQM}} < \Lambda_{B,d} \Rightarrow \text{stable}$$

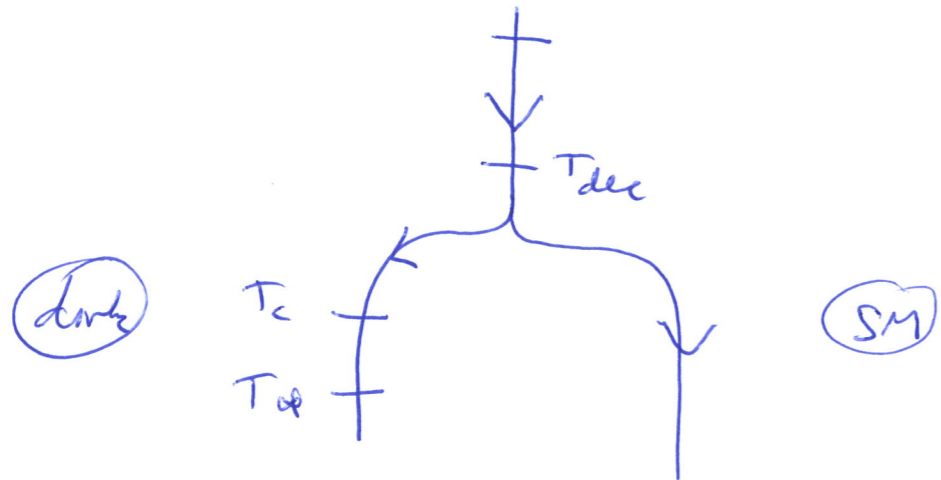
↑
lightest stable dark baryons

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How to make them

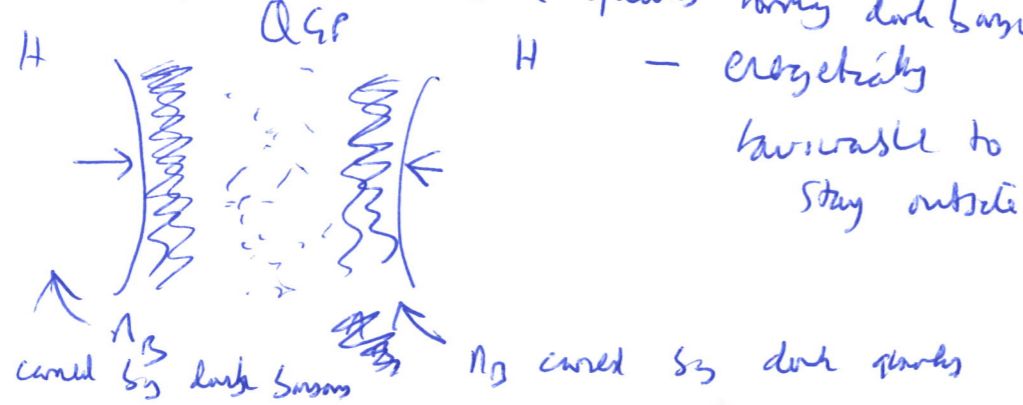
(TL, DR: Same as ordinary quark nuggets)

① Dark sector $\xleftrightarrow{\text{thermal}}$ SM sector
until T_{dec}



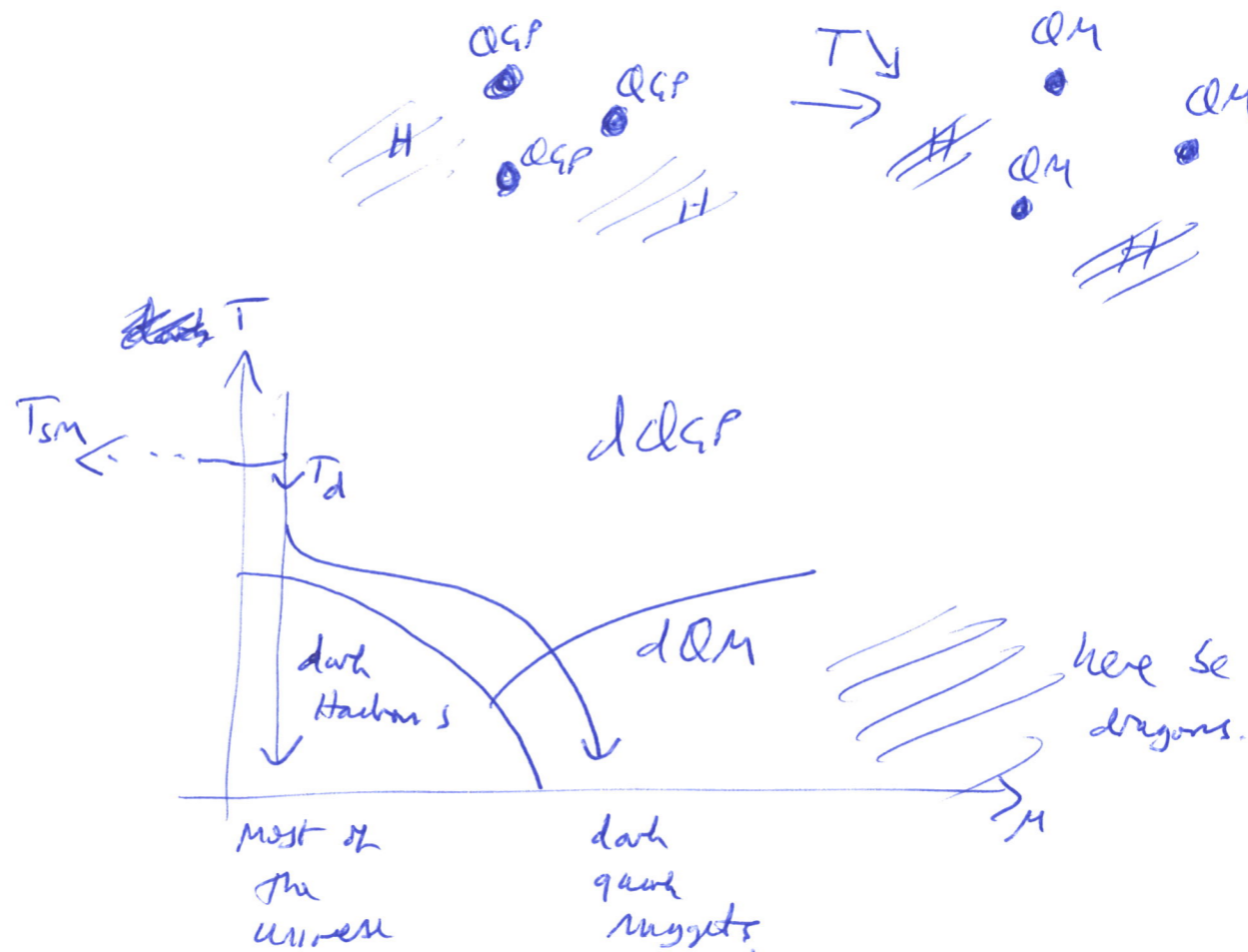
② At $T_q \lesssim T_c$ (Small supercooling)
 [→ not actually sure why this is necessary]
 bubbles of dark hadron phase born
 and grow ~~at~~ at a [non-relativistic] terminal speed

③ Dark ~~quarks~~ ^{quarks} get trapped in front of the walls - because $h \ll m_{SD}$ in the hadron phase is $\gg T$, there is Boltzmann suppression, preventing the dark quarks having dark baryons.



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④ Bubbles of hadron phase coalesce but there are small regions of dark QGP.
 ⑤ These cool to form dark quark nuggets, carrying most of the dark Λ_{SD} .



(7)

Typical nugget size.

- Assuming they don't evaporate
- estimate 1 nugget / 1 hadron bubble
- Use XEFT to ~~est~~ study phase transition
to obtain number of hadron bubbles nucleated.

For $T_d = 100 \text{ MeV}$, $B = 100 \text{ MeV}$

$$\frac{\Lambda_{SU}}{\Lambda_{\gamma}} \approx \frac{\Lambda_B}{\Lambda_{\gamma}}$$

$$M_{\text{dew}} \sim 10^{11} \text{ g}$$

$$R_{\text{dew}} \sim 1 \text{ mm.}$$

Δ the relic abundance can explain all DM